

MIRIS Pa α Observation

1. Observation data
2. Status of data reduction
3. Scientific analysis : Comparison with IPHAS H α data in Cepheus (Q2)

KASI
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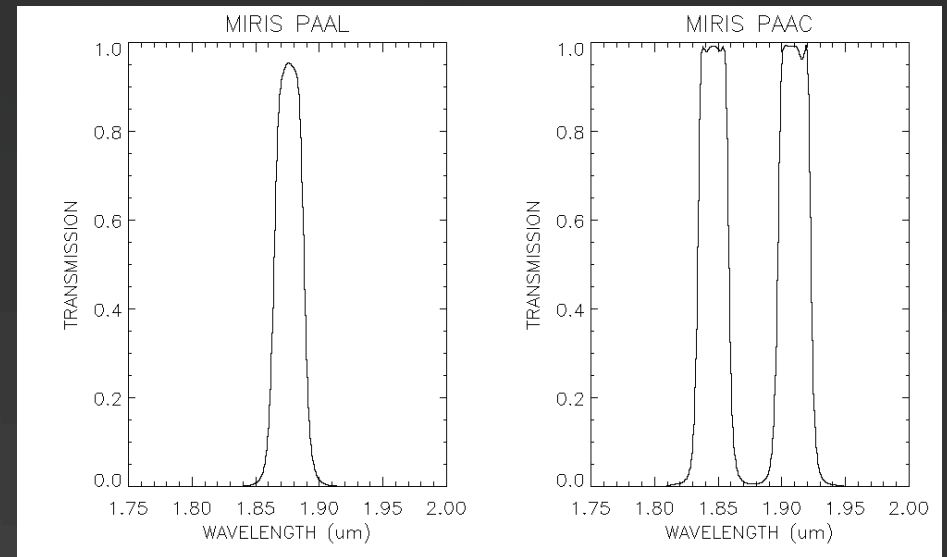
Instrument Specification

◆ MIRIS (Multi-purpose IR Imaging System)

+ The primary payload of the Korean science and technology satellite 3, launched on 2013.

+ Specifications

- Pixel scale : $51.6'' \times 51.6''$
- Field of view : $3.67^\circ \times 3.67^\circ$
- Filters : I ($1.05 \mu\text{m}$), H ($1.6 \mu\text{m}$),
Pa α line ($1.875 \mu\text{m}$),
Pa α dual continuum (1.84 & $1.91 \mu\text{m}$)



+ Main Goals

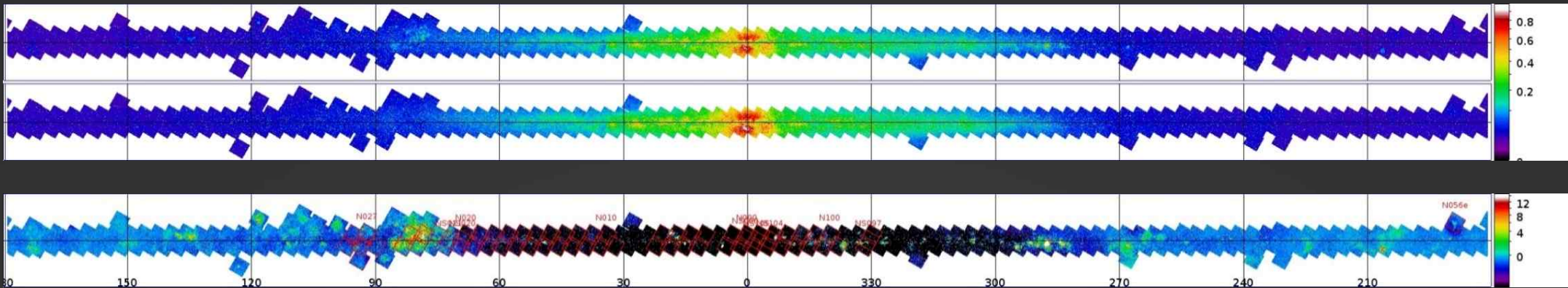
- **Pa α emission line survey of the whole Galactic plane.**
- Cosmic Infrared Background (CIB) observation in NEP & SEP.

1. Observation data

1.1 Observation Data

◆ MIRIS Pa α Galactic Plane Survey (MIPAPS)

- + Cover the whole plane ($3b^{\circ}$) with the width of $-3^{\circ} < b < +3^{\circ}$.
- + Total 235 fields with the average exposure of ~ 20 minutes (per filter).



Top : Pa α line filter (PAAL) image (mJy/arcsec²)

Middle : Pa α dual continuum filter (PAAC) image (mJy/arcsec²)

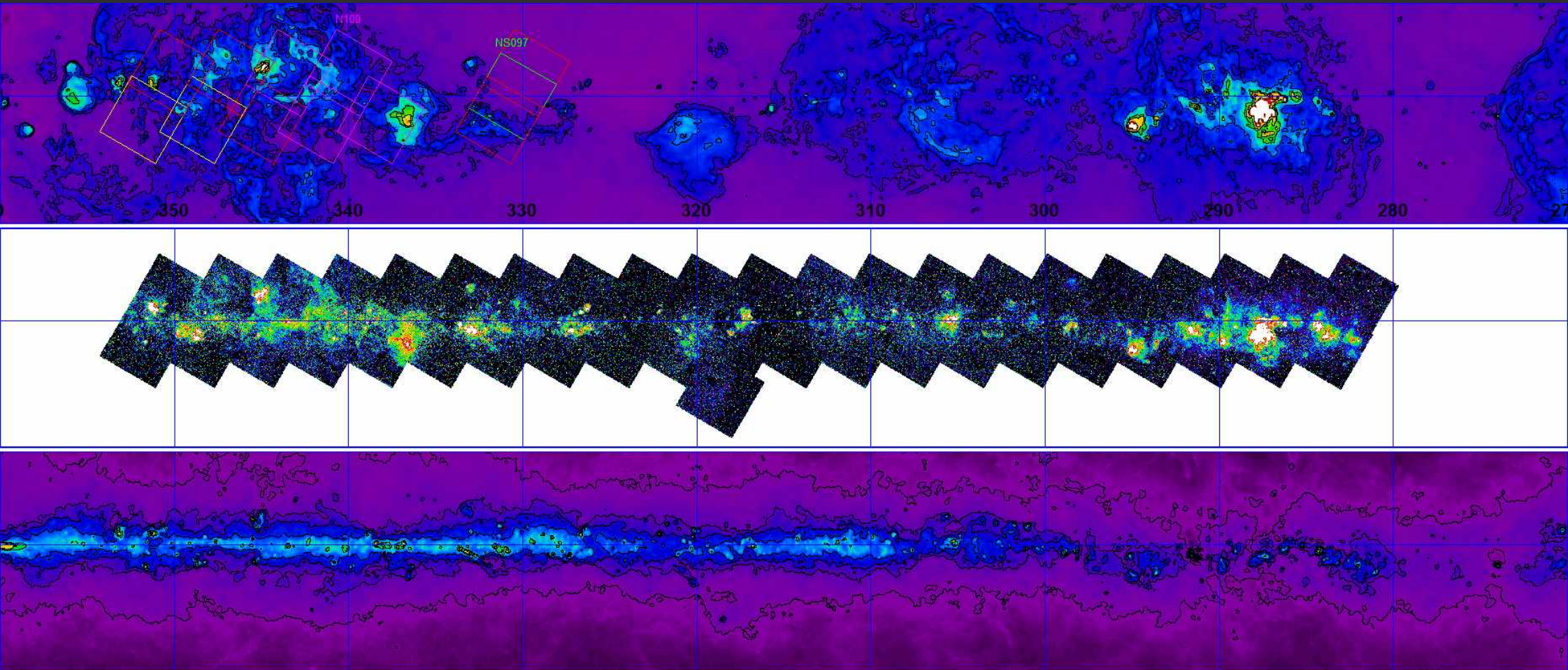
Bottom : Pa α emission line (PAAL-PAAC) image (10^{-19} W/m²/arcsec²) (preliminary)

- + The 1st data release on June, 2017. (<http://miris.kasi.re.kr/miris/>)
- + Need correction of edge-shadowing (by filter-wheel position offset) for $b \approx 2$ orbit data.

1.1

Observation Data

◆ $l = +28^{\circ}$ to $+35^{\circ}$



Top : $H\alpha$ image (Finkbeiner, 2003)

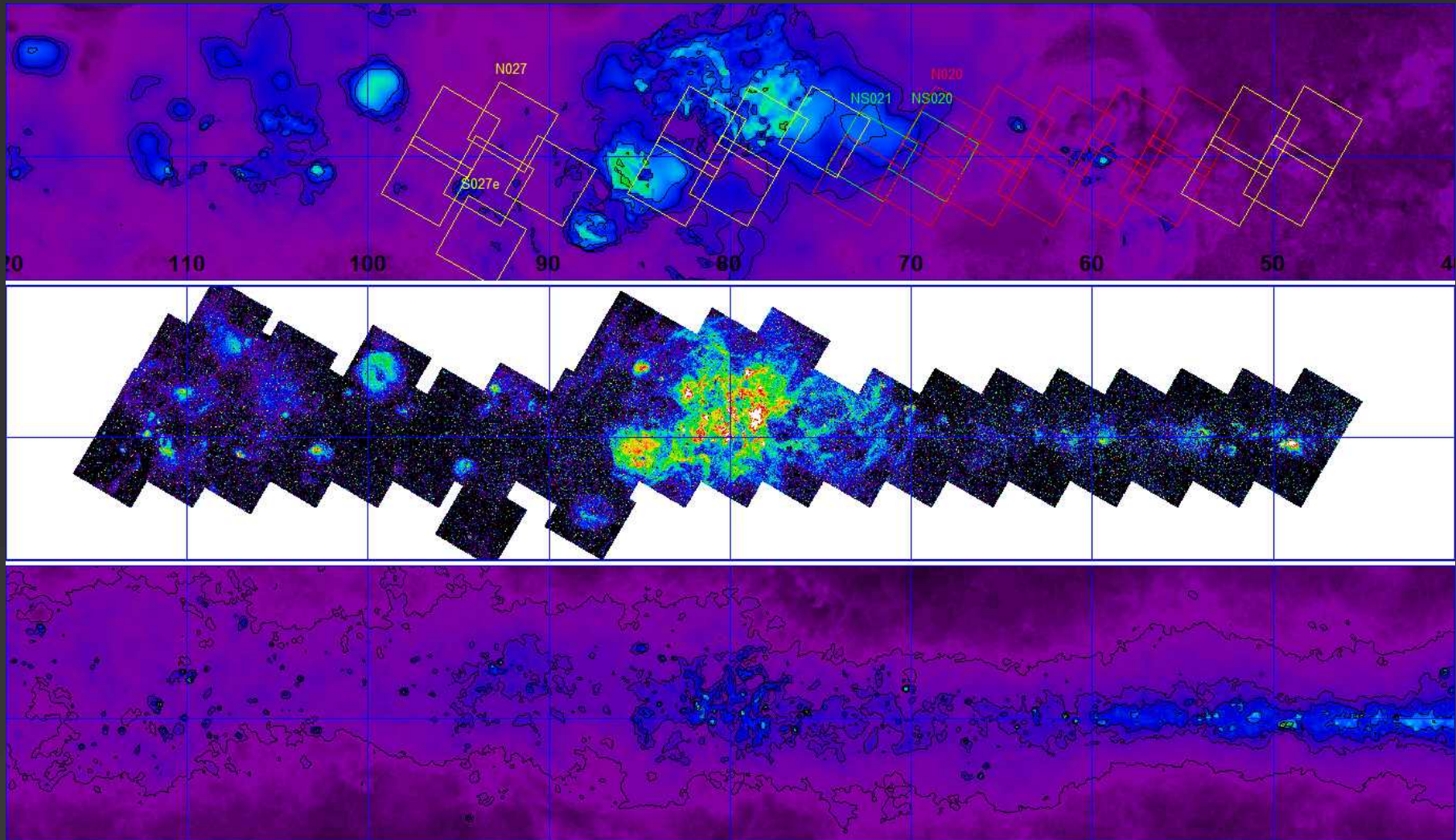
Middle : MIRIS $Pa\alpha$ image

Bottom : SFD $E(B-V)$ dust image (Schlegel+ 1998)

1.1

Observation Data

◆ $l = +050^\circ$ to $+110^\circ$



Top : $H\alpha$ image (Finkbeiner, 2003)

Middle : MIRIS $Pa\alpha$ image

Bottom : SFD E(B-V) dust image (Schlegel+ 1998)

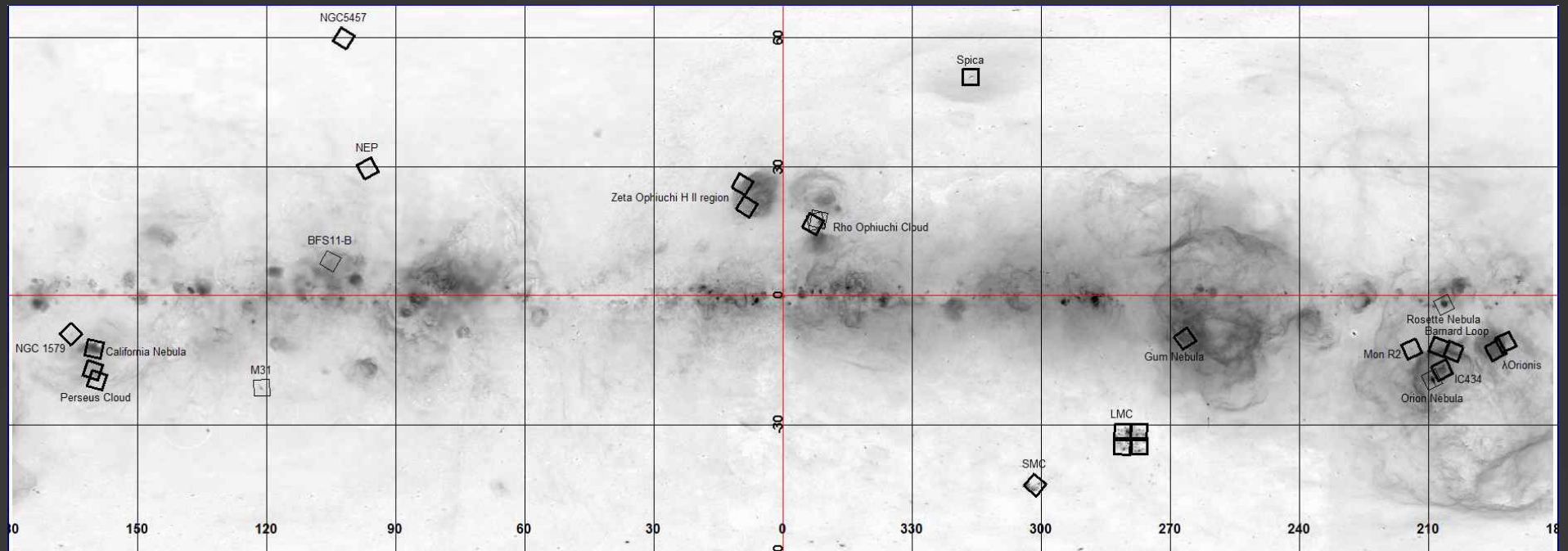
1.2

Observation Data

◆ Pointing observations

+ Cover **19 targets** (total 26 fields) located away from the Galactic plane.

- Nearby H II regions : M42, Rosette nebula, Lambda Orionis, California nebula, IC434, Barnard Loop, Gum nebula, Spica nebula, Zeta Ophiuchus.
- Star-forming clouds : Rho Ophiuchus, BFS11-B, Perseus cloud, NGC1579, Mon R2 cloud.
- Nearby galaxies : M31, SMC, LMC, NGC5457.
- North eclipse pole (NEP).



Positions of Pa α pointing observations on H α all sky map

2. *Status of data reduction*

◆ Flux calibration

- + Point sources, Extended sources, Check PAAL/PAAC PSF radial profile.
- + Check detector-counting variation.

◆ Edge-shadowing by filter-wheel position offset

- + Check the status of affected data.
- + Crop or correct for b02 orbit data by stray light simulations.

◆ Background increment & contamination features by stray light

- + Check the degree of background increment by Moon.
- + Check and correct on individual data analysis.

◆ Stellar continuum subtraction

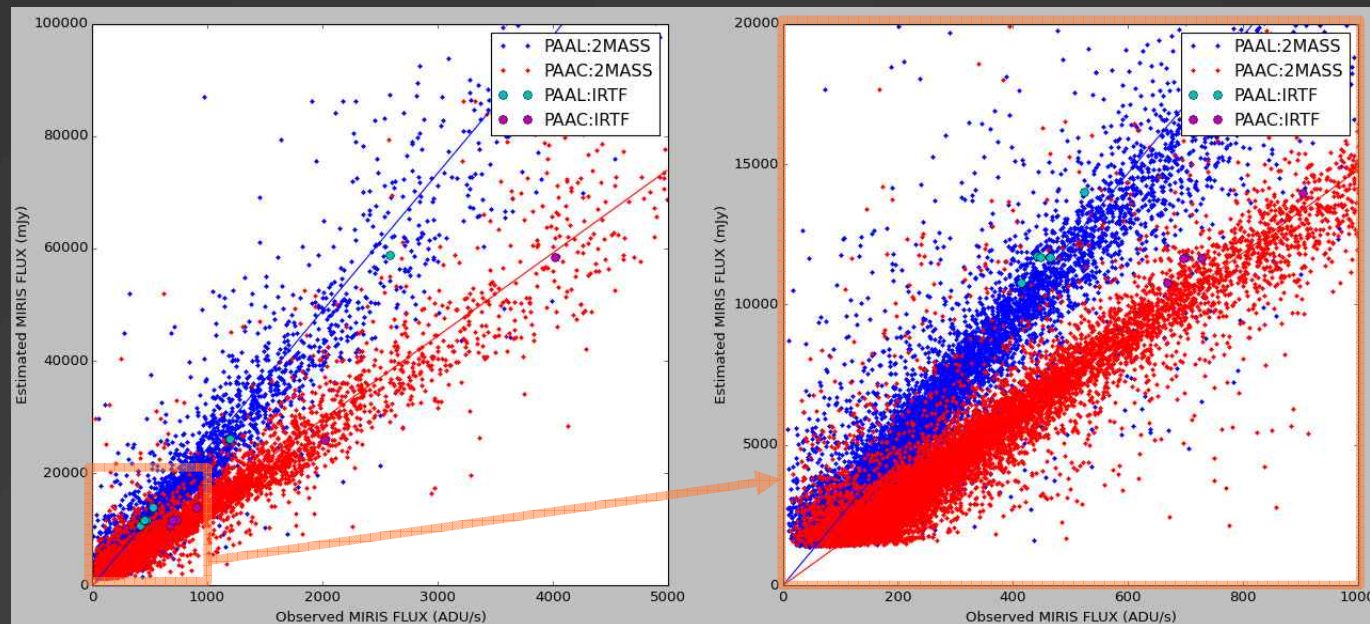
- + PSF estimation (by StarFinder) and PSF matching (by Photutils).
- + Mask and fill by using point source catalog.
- + Find more effective method.
- + For the whole plane image.

2.1

Status of data reduction

◆ Flux calibration

- + MIRIS sources : extraction and aperture photometry (**ADU/sec**) by SExtractor.
- + 2MASS sources : $H < 7$ & $K < 7$ from 2MASS point source catalog.
 - Matching sources : **16676** (PAAL), **15344** (PAAC) from the whole Galactic plane.
- + Estimate **mJy** flux at MIRIS PAAL/PAAC from **2MASS H & K** magnitude (interpolation).
- + Obtain **calibration factor** from **ADU/sec** to **mJy** by linear fitting.
 - PAAL : **24.5**, PAAC : **14.8**.

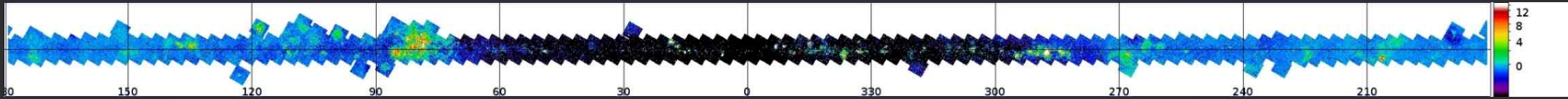


Observed flux (ADU/sec) vs. Estimated flux (mJy)

2.1

Status of data reduction

◆ Flux calibration



Pa α emission line (PAAL-PAAC) image (10^{-19} W/m²/arcsec²) (preliminary)

+ Effects by detector-counting variation or unknown factors.

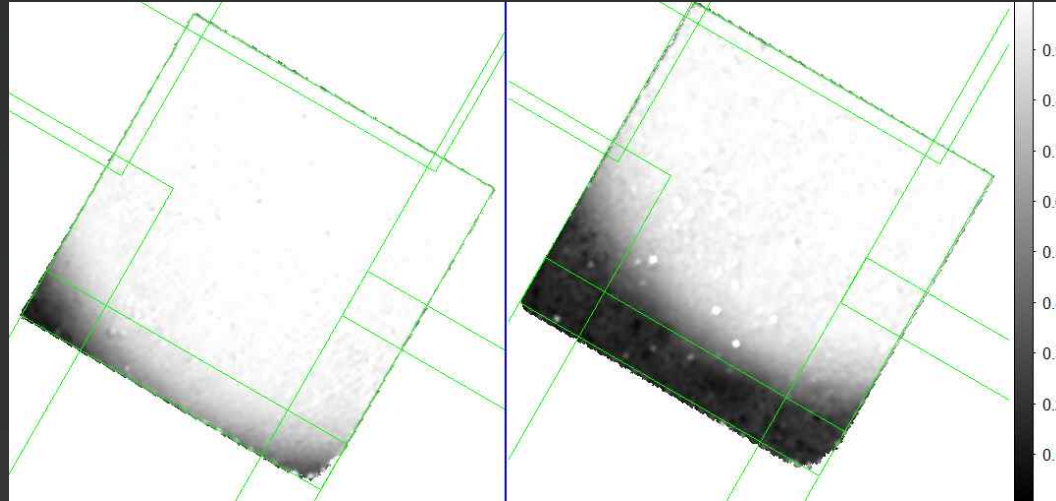
- Why is PAAC background much larger than PAAL background around GC?
- Why is MIRIS ADUs more under-estimated for brighter pixels, PAAC?

+ Need to work

- Find the cause and correct.

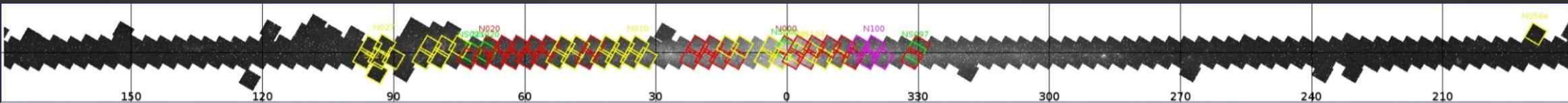
Status of data reduction

◆ Edge-shadowing by filter-wheel position offset



+ Check the status of affected data.

- MIPAPS : total b02 orbit data (~30%) were affected.
 - >50% dimmed (143 orbits), 20-50% dimmed (203 orbits), <20% dimmed (256 orbits).



- **Pointing observations : total 3 target data were affected.**

<45% dimmed (Rho Ophiuchus), <20% dimmed (LMC, Zeta Ophiuchus)

+ Need to work

- Crop or correct for b⁰₂ orbit data by stray light simulations.

3. Scientific analysis : Comparison with IPHAS H_{α} data in Cepheus (Q2)

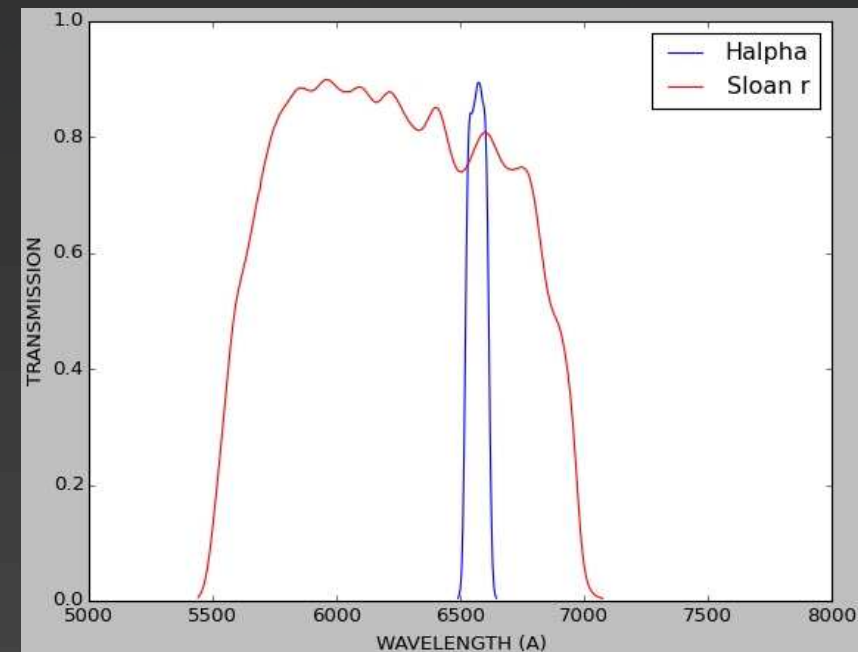
3.1

Data : IPHAS $H\alpha$

◆ Isaac Newton Telescope (INT) Photometric $H\alpha$ Survey (IPHAS)

+ The 2nd data release : the first quality-controlled and globally calibrated data (Barentsen+ 2014).

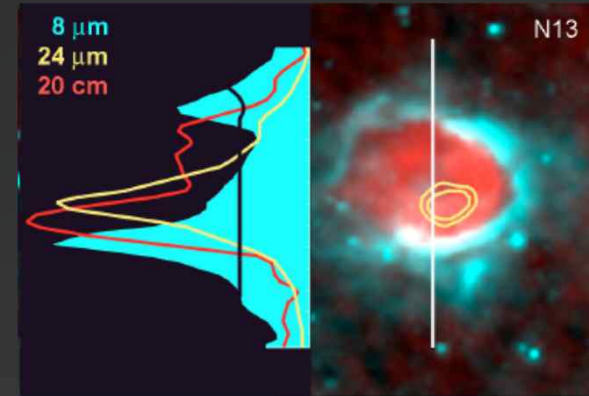
Property	Value
Telescope	2.5-m INT
Instrument	Wide Field Camera (WFC)
Detectors	Four 2048 × 4100 pixel CCDs
Pixel scale	0.33 arcsec pixel ⁻¹
Filters	r , i , $H\alpha$
Filter properties	See Table 2
Magnitude system	Vega
Exposure times	30 s (r), 10 s (i), 120 s ($H\alpha$)
Saturation limit	13 (r), 12 (i), 12.5 ($H\alpha$)
Detection limit (5σ , mean)	21.2 (r), 20.0 (i), 20.3 ($H\alpha$)
PSF FWHM (median)	1.1 arcsec (r), 1.0 arcsec (i), 1.1 arcsec ($H\alpha$)
Survey area	~1860 deg ²
Footprint boundaries	$-5^\circ < b < +5^\circ$, $29^\circ < \ell < 215^\circ$
Observing period	2003 August–2012 November
Website	www.iphas.org



3.1 Data : WISE H II region catalog

◆ Previous observations of H II regions

- + Ionized Hydrogen gas : H_α , radio recombination line (RRL), radio continuum.
- + PAHs and/or VSGs : $\sim 10\ \mu\text{m}$ MIR.
- + VSGs or BGs : $\sim 20\ \mu\text{m}$ MIR.



Deharveng+ 2010

◆ WISE catalog of Galactic H II regions (Anderson+ 2014)

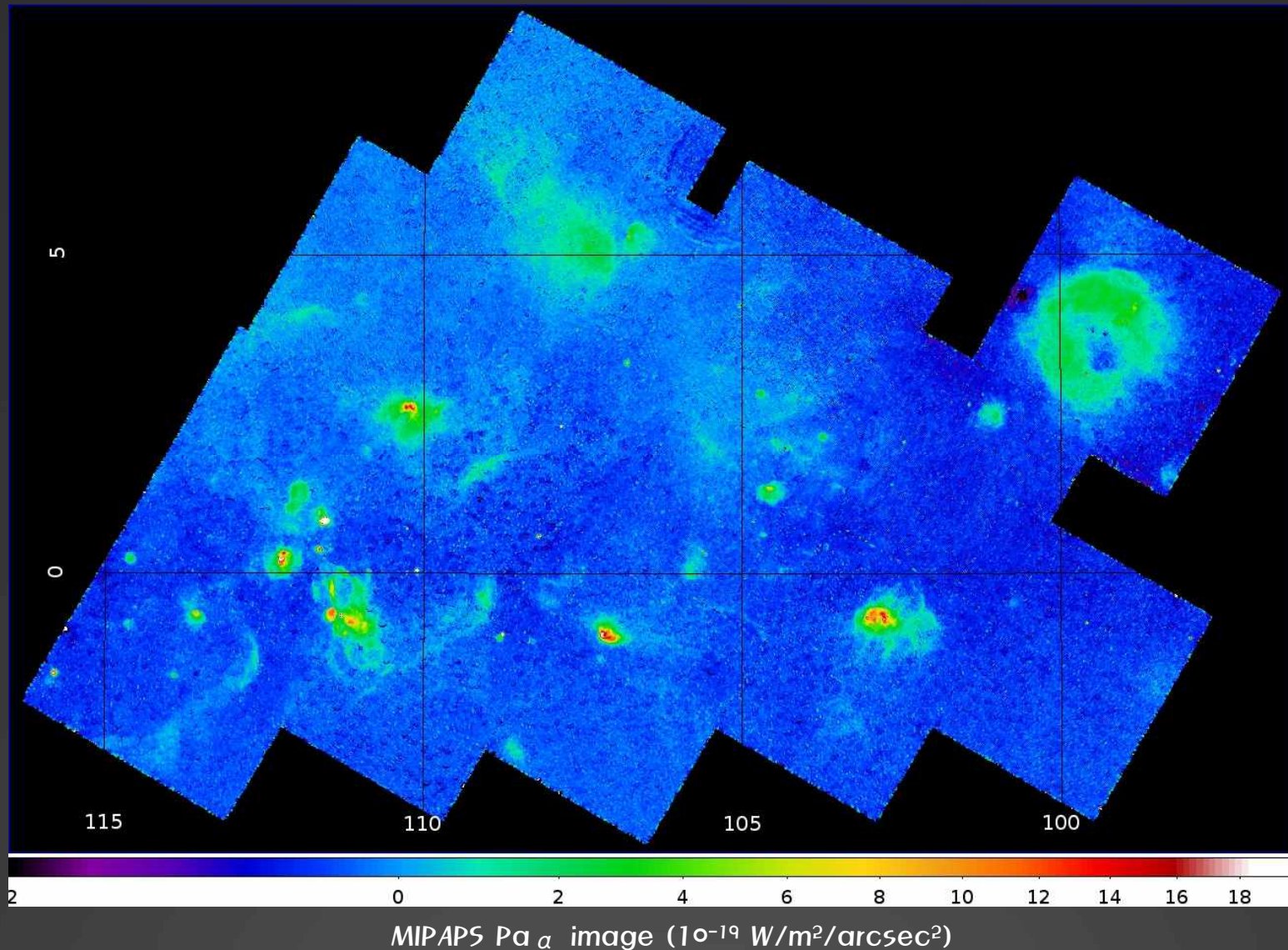
- + WISE has sufficient sensitivity (6 mJy at $22\ \mu\text{m}$) to detect the MIR emission from H II regions located anywhere in the Galactic disk.
- + The most complete catalog : total 8399 sources (WISE $12\ \mu\text{m}$ + $22\ \mu\text{m}$)
 - Known (1524) : hydrogen recombination line (H_α or RRL) detected.
 - Candidate (1986) : radio continuum detected, but no H_α or RRL detected.
 - Group (650) : overlapped with other Known sources.
 - Radio quiet (4124) : no radio, H_α or RRL detected (only $10\ \mu\text{m}$ & $20\ \mu\text{m}$ MIR detected).

3.2

Mosaic images in Cepheus

◆ Continuum-subtracted Pa_α mosaic images from MIPAPS

+ Use 14 fields data (PAAL, PAAC filters) : flux-calibrated by comparing 2MASS point source catalog.

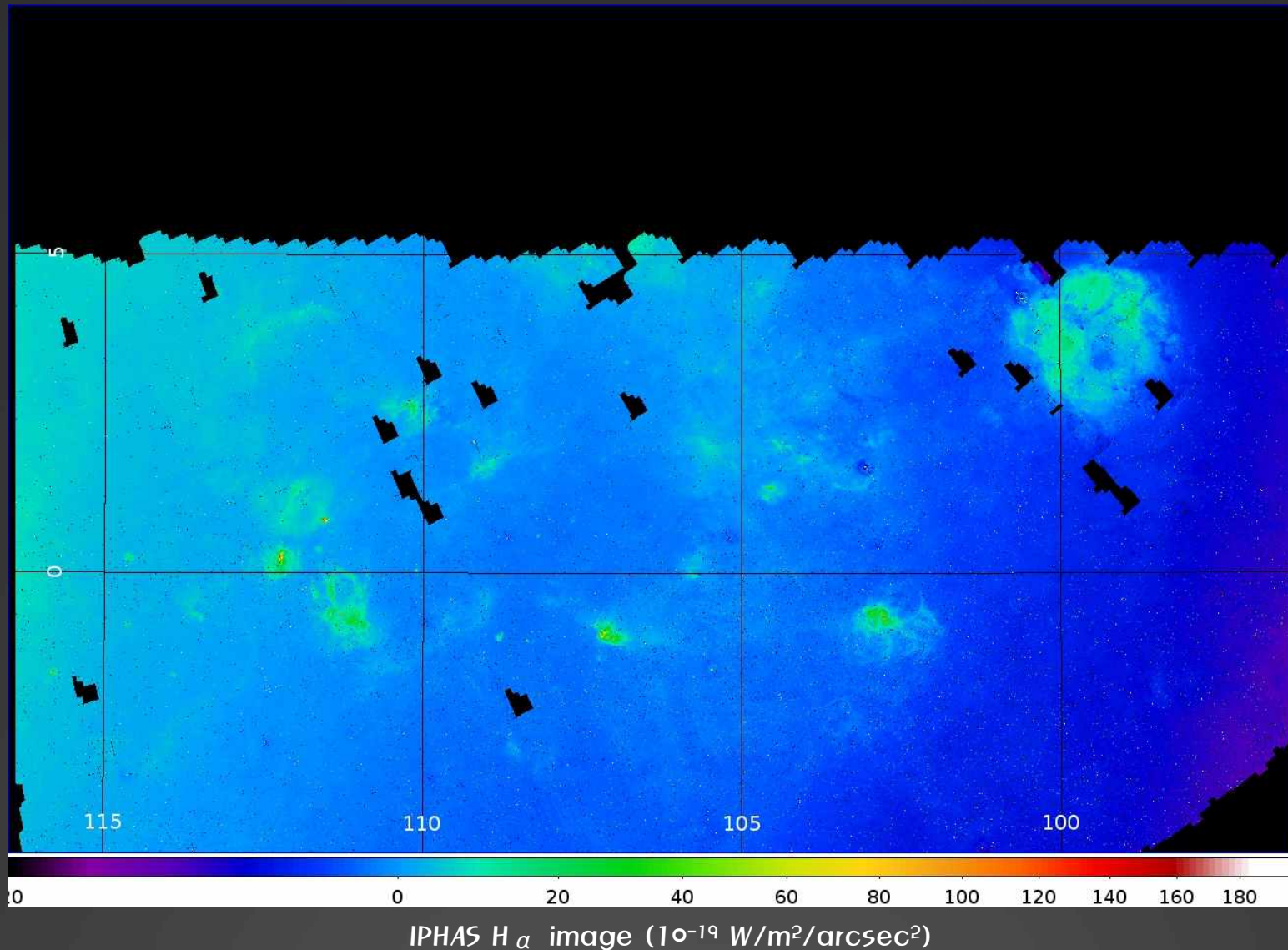


3.2

Mosaic images in Cepheus

◆ Continuum-subtracted H_α mosaic images from *IPHAS*

+ Use 7840 fields data (halpha, r filters) : 15 pixel binning (pixel size = 4.9 arcsec).

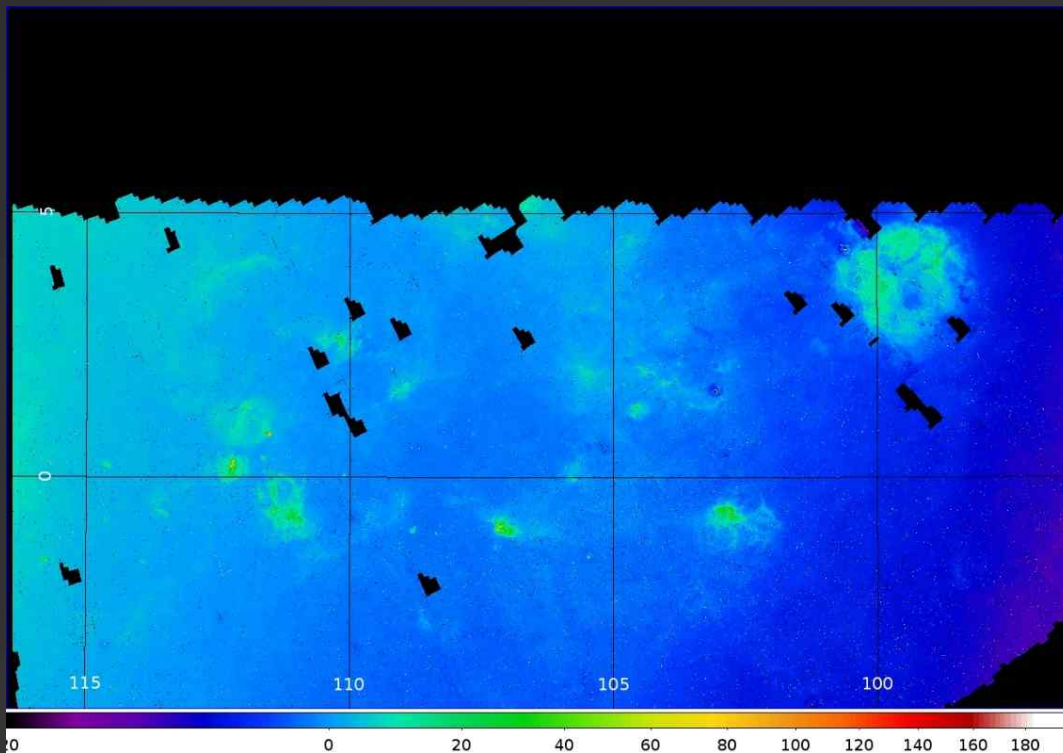


3.2

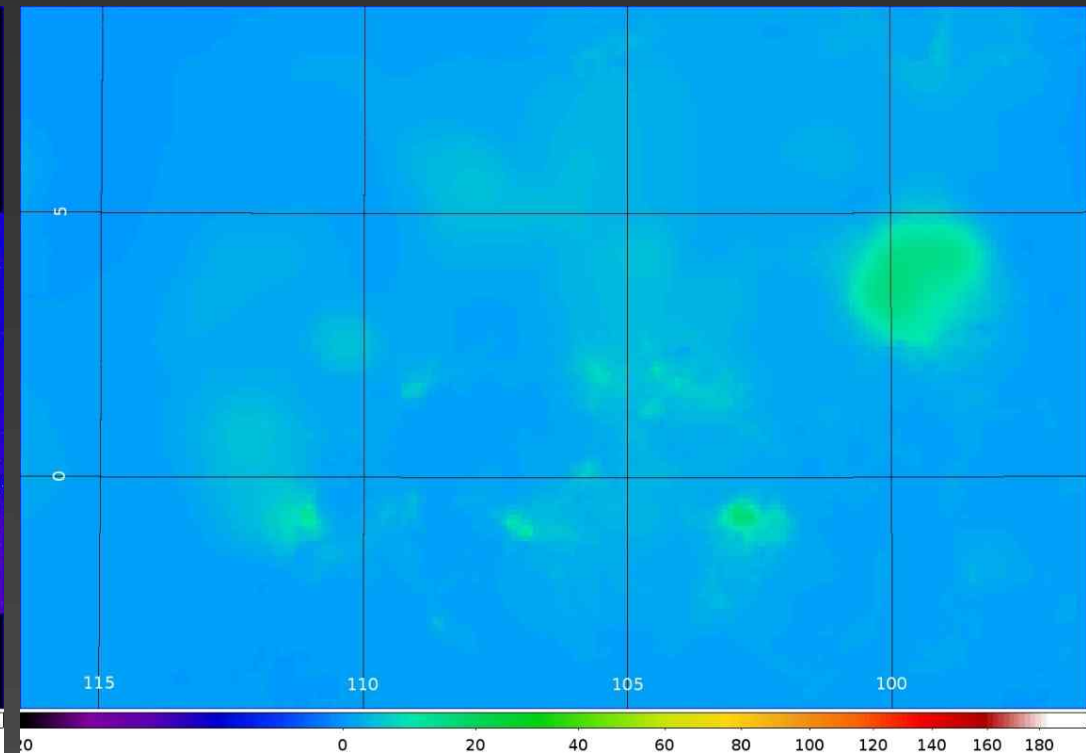
Mosaic images in Cepheus

◆ Continuum-subtracted H_α mosaic images from IPHAS

+ Comparison with the previous H_α image (Finkbeiner 2003).



IPHAS H_α image
($10^{-19} \text{ W/m}^2/\text{arcsec}^2$)

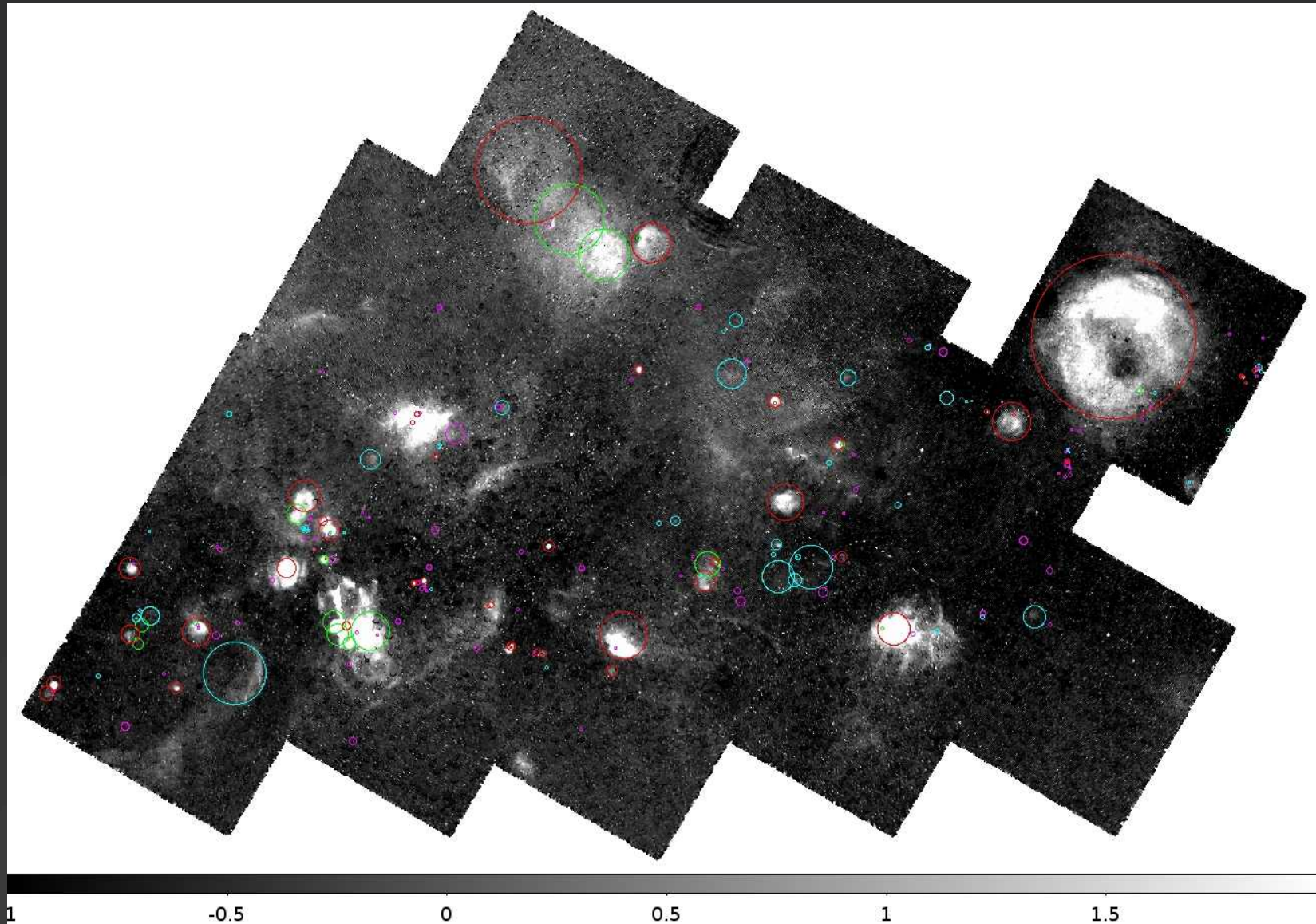


Finkbeiner's H_α image
($10^{-19} \text{ W/m}^2/\text{arcsec}^2$)

3.2

WISE H II region sources

◆ 49 **Known**, 50 **Candidate**, 21 **Group**, 92 **Radio quiet**



MIPAP5 Pa α image (10^{-19} W/m²/arcsec²)

3.2

WISE H II region sources

◆ Results of visual inspection & flux photometry

Name	Positions	Radius	Paa feature	Overlapped with PS	IPHAS	Paa Int	Paa Flux	Ha Int	Ha Flux	E(B-V)	Dist.	Lyc Flux	SpType
WK01	G097.311+03.269	116	Y		Y	0.79	0.34	0.18	0.07	1.93	9200	47.68	O9.3V
WK02	G097.444+03.083	95	Y		Y	0.27	0.11	0.84	0.34	0.53	10400	47.04	>O9.5V
WK03	G097.515+03.173	122	Oy		Y	4.73	2.56	2.79	1.48	1.40	7500	48.29	O8.0V
WK04	G097.528+03.184	68	On		Y								
WK05	G099.484+03.801	4539	Y	partially	Y	1.84	1109.90	11.86	7032.08	0.15	860	48.81	O6.5V
WK06	G100.205+01.885	166	Oy		Oy	0.22	0.12						
WK07	G101.016+02.590	101	On		Y								
WK08	G101.065+02.499	1066	Oy	partially	Y	1.16	31.88	1.05	28.84	1.17	7300	49.32	O4.7V
WK09	G101.439+02.653	136	Y		Y	1.77	1.05	1.02	0.61	1.41	7200	47.86	O9.1V
WK10	G102.877-00.695	913	Oy	partially	Y	5.51	164.49	14.34	430.46	0.62	5800	49.73	<O3.0V
WK11	G103.686+00.425	290	Oy	partially	Y	0.49	0.91	0.12	0.22	1.87	7400	47.91	O9.0V
WK12	G103.743+02.162	342	Oy		Y	1.33	4.27	1.84	6.00	0.94	2700	47.54	>O9.5V
WK13	G104.546+01.255	1007	Y	partially	Y	1.11	30.99	2.85	79.48	0.63	3100	48.46	O7.4V
WK14	G104.700+02.784	102	On		Y								
WK15	G104.716+02.813	353	Oy	partially	Y	1.29	4.39	0.44	1.51	1.67	7800	48.60	O7.1V
WK16	G105.635+00.345	227	Oy	largely	Y								
WK17	G105.779+00.048	584	Oy	partially	Y	1.01	10.84	3.82	40.87	0.43	5200	48.42	O7.6V
WK18	G106.605+05.252	1107	Oy	partially	partial, Y	0.91	36.25						
WK19	G106.809+03.310	256	Y		Y	1.40	2.94	1.20	2.46	1.21	7000	48.25	O8.1V
WK20	G107.034-00.801	1323	Oy	partially	Y	2.31	127.25	8.26	454.64	0.45	4300	49.33	O4.7V
WK21	G107.209-01.334	326	Y		Y	0.47	1.29	2.03	5.60	0.35	4100	47.27	>O9.5V
WK22	G108.191+00.586	307	Y		Y	9.14	10.13	2.65	2.92	1.77	4400	48.49	O7.4V
WK23	G108.273-01.066	195	Y	largely	Y								
WK24	G108.375-01.056	187	Oy	largely	Y								
WK25	G108.503+06.356	2940	Oy	partially	no data	0.43	116.98						
WK26	G108.752-00.972	42	On		N								
WK27	G108.758-00.989	46	On		Oy								
WK28	G108.764-00.952	237	Oy		Y	11.38	9.85	14.50	13.24	0.96	5600	48.54	O7.2V
WK29	G108.770-00.974	38	On		Oy								
WK30	G109.068-00.322	180	Oy		Oyw	0.85	0.85						
WK31	G109.104-00.347	95	Oy		Y	2.19	0.71						
WK32	G109.927+01.981	179	Y		Y	0.29	0.17	0.64	0.35	0.74	5000	46.64	>O9.5V
WK33	G110.099+00.042	192	Y		Y	11.62	15.71	11.99	16.30	1.10	5500	48.75	O6.6V
WK34	G110.211+02.616	140	Oy	partially	Oy	6.17	3.84						
WK35	G110.252+00.009	153	Y	partially	Y	1.07	0.40	1.10	0.44	1.07	6400	47.29	>O9.5V
WK36	G110.286+02.488	117	On		N								
WK37	G111.286-00.660	236	Y		Y	7.85	4.88	19.75	13.05	0.60	3400	47.73	O9.2V
WK38	G111.558+00.804	573	Oy	partially	Y	30.81	126.62	17.78	73.53	1.40	2600	49.06	O5.7V
WK39	G111.612+00.371	54	Oy		Y	18.11	1.96	7.65	0.65	1.69	7200	48.19	O8.3V
WK40	G111.653+00.950	273	Oy	partially	Oyw	2.24	5.22						
WK41	G111.802+00.526	96	Y		Y	0.41	0.14	0.10	0.04	1.86	5000	46.77	>O9.5V
WK42	G111.946+01.336	880	Oy	partially	Y	1.01	24.43	2.40	58.01	0.67	2800	48.28	O8.0V
WK43	G112.212+00.229	542	Y	partially	Y	6.10	56.25	22.43	206.86	0.44	4600	49.03	O5.8V
WK44	G113.595-00.749	783	Oy	partially	Y	1.61	23.44	1.39	20.41	1.19	4400	48.75	O6.7V
WK45	G113.900-01.613	327	Y		Y	0.54	2.82	1.06	5.66	0.75	4800	47.82	O9.1V
WK46	G114.605-00.801	522	Oy	partially	Y	0.44	3.67	0.68	5.77	0.88	3400	47.66	O9.4V
WK47	G114.626+00.219	593	Y	partially	Y	1.40	6.25	3.85	17.16	0.59	4900	48.16	O8.4V
WK48	G115.785-01.561	402	Y		Y	3.89	12.82	6.96	23.26	0.81	4400	48.41	O7.6V
WK49	G115.885-01.707	406	Y		Y	0.51	1.42	0.94	2.59	0.80	4500	47.48	>O9.5V

49 Known sources

radius (arcsec), intensity (10^{-19} W/m²/arcsec²), flux (10^{-14} W/m²)

Name	Positions	Radius	Paa feature	Overlapped with PS	IPHAS	Paa Int	Paa Flux	Ha Int	Ha Flux	E(B-V)	Dist.	Lyc Flux	SpType
WC01	G097.210+03.245	100	N		N								
WC02	G097.252+03.320	189	Oy		YWW	0.56	0.64	0.10	0.11	2.01	5000	47.44	>O9.5V
WC03	G097.728+02.352	108	On		Y								
WC04	G098.320+01.552	72	On	partially	On								
WC05	G098.855+02.933	91	Oyw		Y	0.16	0.04						
WC06	G100.169+02.026	44	N	totally	N								
WC07	G100.181+02.038	54	N	totally	Y								
WC08	G100.199+02.064	81	Y	largely	Y								
WC09	G100.714-00.527	638	Y	partially	Y	0.25	3.24	0.32	4.04	1.00	5000	47.96	O8.8V
WC10	G101.527-00.515	116	N	partially	Y								
WC11	G101.663+02.820	55	N		N								
WC12	G101.763+02.808	77	N		N								
WC13	G102.051+02.861	380	N	partially	N								
WC14	G102.207-00.736	134	On		Y								
WC15	G102.327+03.681	40	N		Y								
WC16	G102.354+03.635	149	N		Y								
WC17	G102.807+01.204	186	Y		Y	0.19	0.20	0.04	0.04	1.92	5000	46.91	>O9.5V
WC18	G103.578+03.165	445	Y	partially	Y	0.15	0.92	0.21	1.30	0.94	5000	47.41	>O9.5V
WC19	G103.875+01.857	164	On		On								
WC20	G104.153+00.259	1220	Oy	partially	Y								
WC21	G104.355+00.404	157	On		On								
WC22	G104.393+00.049	375	Oy		partial, Y	0.37	0.78	0.62	1.32	0.85	5000	47.32	>O9.5V
WC23	G104.648+00.110	907	Oy	partially	Y								
WC24	G104.675+00.595	305	Y		Y	0.65	1.45	0.40	0.88	1.38	5000	47.68	O9.3V
WC25	G104.728+00.446	126	Oyw	partially	N	0.09	0.04						
WC26	G105.307+04.058	363	On	partially	N								
WC27	G105.365+03.234	829	Oy	partially	Y	0.15	3.20	0.80	17.16	0.24	5000	47.82	O9.1V
WC28	G105.487+03.889	109	On		N								
WC29	G105.852+00.142	129	Oy		Y	0.27	0.14	0.05	0.03	2.01	5100	46.81	>O9.5V
WC30	G106.241+00.957	274	Y	partially	YWW	0.09	0.20						
WC31	G106.499+00.925	123	N		N								
WC32	G108.213-01.293	103	YW		YWW	0.17	0.06						
WC33	G108.902+02.714	410	Oy	partially	partial, Y	0.11	0.59						
WC34	G109.854+02.147	85	On	totally	N								
WC35	G109.874+02.115	129	Oy	partially	YW	1.47	0.75	0.43	0.23	1.75	700	45.76	>O9.5V
WC36	G109.994-00.092	92	N		Y								
WC37	G110.923+01.906	566	Y	partially	Y	0.33	2.92	0.51	4.53	0.89	5000	47.90	O9.0V
WC38	G111.329+00.783	60	On		On								
WC39	G111.873+00.820	60	On	totally	On								
WC40	G111.907+00.800	99	On		On								
WC41	G111.922+00.859	113	On		On								
WC42	G111.947+00.799	104	On		On								
WC43	G111.966+00.839	113	Oy	partially	Oyw	0.47	0.20	0.37	0.15	1.29	5000	46.82	>O9.5V
WC44	G113.009-01.393	1734	Y	partially	Y	0.19	21.36						
WC45	G113.096+02.602	159	Oyw		N	0.09	0.08						
WC46	G114.312-00.510	501	Oy	partially	Oy	0.18	1.40	0.38	2.99	0.72	5000	47.55	>O9.5V
WC47	G114.332+00.788	60	N		N								
WC48	G114.473-00.430	88	Y		YW	0.011	0.003	0.010	0.002	1.15	5000	44.89	>O9.5V
WC49	G114.526-00.543	229	Y	partially	YW	0.23	0.38	0.05	0.08	1.93	5000	47.20	>O9.5V
WC50	G115.105-01.438	124	Y		YW	0.26	0.13						

50 Candidate sources

3.2

WISE H II region sources

◆ Results of visual inspection & flux photometry

Name	Positions	Radius	Paa feature	Overlapped with PS	IPHAS	Paa Int	Paa Flux	Ha Int	Ha Flux	E(B-V)	Dist.	Lyc Flux	SpType
WG01	G099.091+02.969	118	Oy		Oy	0.72	0.31	2.29	1.00	0.51	5000	46.86	>O9.5V
WG02	G103.061-00.691	104	Oy		Oy	3.91	1.48	3.55	1.21	1.23	5000	47.66	O9.3V
WG03	G103.659+02.151	120	Oyw		Oyw	0.36	0.12	0.45	0.15	0.98	5000	46.52	>O9.5V
WG04	G105.571+00.296	139	Oy		Oy	0.74	0.44	0.66	0.40	1.16	5000	47.13	>O9.5V
WG05	G105.744+00.298	700	Oy	partially	Oy	0.71	10.98	2.15	33.13	0.54	5000	48.41	O7.6V
WG06	G106.798+05.313	135	Oy	partially	no data	1.49	0.85						
WG07	G107.333+05.061	1440	Oy	partially	partial, Y	1.29	84.61						
WG08	G107.866+05.607	1968	Oy	partially	no data	0.75	91.51						
WG09	G110.927-00.731	1058	Oy	partially	Oy	2.84	100.20	7.58	266.59	0.61	5000	49.38	O4.4V
WG10	G111.196-00.798	73	On		On								
WG11	G111.245-00.924	345	Oy		Oy	0.62	2.34	1.34	4.99	0.72	5000	47.77	O9.2V
WG12	G111.430-00.790	587	Oy	partially	Oy	3.04	33.01	5.12	55.42	0.85	5000	48.94	O6.1V
WG13	G111.478-00.591	620	Oy	partially	Oy	2.91	34.96	5.42	65.36	0.79	5000	48.96	O6.0V
WG14	G111.543+00.775	69	On	largely	OyW								
WG15	G111.601+00.393	62	On		Oy								
WG16	G111.640+00.360	232	Oy		Oy	6.91	11.77	4.95	8.37	1.30	5000	48.58	O7.1V
WG17	G112.071+01.063	533	Oy	partially	Oy	1.30	11.53	1.98	17.66	0.90	5000	48.50	O7.4V
WG18	G113.614-00.615	112	On		Oy								
WG19	G114.437-00.662	381	Oy	partially	Oyw	0.07	0.31						
WG20	G114.496-00.946	282	Oy	partially	Oyw	0.11	0.28	0.30	0.76	0.60	5000	46.83	>O9.5V
WG21	G114.520-00.867	60	On		Oy								

Name	Positions	Radius	Paa feature	Overlapped with PS	IPHAS	Paa Int	Paa Flux	Ha Int	Ha Flux	E(B-V)	Dist.	Lyc Flux	SpType
WR37	G105.227-00.262	262	N	partially	N								
WR38	G105.279-00.113	193	N	partially	Y								
WR39	G105.509+00.230	99	On		N								
WR40	G105.880+04.253	153	On		N								
WR41	G105.962+00.420	70	N		N								
WR42	G106.142+00.129	75	N		N								
WR43	G106.909+03.148	107	N		YWW								
WR44	G107.156-00.988	69	On		On								
WR45	G107.298+05.638	73	On	partially	no data								
WR46	G107.678+00.235	156	Oyw		N	0.07	0.05						
WR47	G107.683-02.239	100	YW		Y	0.03	0.01	0.015	0.005	1.38	6100	45.57	>O9.5V
WR48	G108.184+05.518	88	On		no data								
WR49	G108.394-01.046	34	N	totally	Oy								
WR50	G108.412-01.097	100	N	partially	N								
WR51	G108.603+00.494	138	N		N								
WR52	G108.666-00.391	93	N		N								
WR53	G108.966+02.726	141	On		no data								
WR54	G109.285-00.987	177	On		N								
WR55	G109.621+02.312	533	On	partially	Oy								
WR56	G109.878+04.261	159	N		N								
WR57	G109.919+00.813	221	N		no data								
WR58	G110.016+00.259	150	N		N								
WR59	G110.054-00.107	60	N		N								
WR60	G110.094-00.064	74	N		Y								
WR61	G110.135-00.077	118	N		N								
WR62	G110.160+00.040	50	N		On								
WR63	G110.170+02.630	106	On	totally	On								
WR64	G110.199+00.016	121	N	totally	YW								
WR65	G110.505-00.586	147	On		N								
WR66	G110.548+02.622	104	On	totally	Y								
WR67	G110.812-00.799	79	On		On								
WR68	G110.941+01.018	83	N		N								
WR69	G111.046+01.085	73	N		N								
WR70	G111.125-00.757	88	On		On								
WR71	G111.180-02.419	204	N		YW								
WR72	G111.236-01.238	132	Oy		Y	1.04	0.59	0.78	0.43	1.29	3500	46.96	>O9.5V
WR73	G111.498+00.369	210	Oyw	partially	Y	0.58	0.81	0.35	0.49	1.38	5000	47.43	>O9.5V
WR74	G111.567+00.751	60	On		On								
WR75	G111.670+03.264	89	N		N								
WR76	G111.774+00.689	75	On		On								
WR77	G111.860+00.800	50	On	largely	On								
WR78	G111.861+01.001	89	Oyw		Oyw	-0.06	-0.01	-0.26	-0.06				
WR79	G111.870+00.881	46	On		On								
WR80	G111.893+00.991	87	Oyw		Oyw	0.07	0.02	-0.22	-0.05		5100		
WR81	G111.941+00.677	75	On		On								
WR82	G112.434+00.034	124	On		On								
WR83	G112.970-00.608	114	On		Oyw								
WR84	G113.246+00.511	137	On		Oy								
WR85	G113.284+00.558	93	On		Oy								
WR86	G113.289-00.819	217	N	partially	N								
WR87	G113.566-00.698	77	On		Oyw								
WR88	G113.569-00.657	77	On		Oyw								
WR89	G114.082-01.401	105	N	totally	Y								
WR90	G114.569+00.290	58	On	partially	On								
WR91	G114.677-02.208	227	On		On								
WR92	G115.782-01.689	90	On	partially	N								

21 Group sources

92 Radio quiet sources

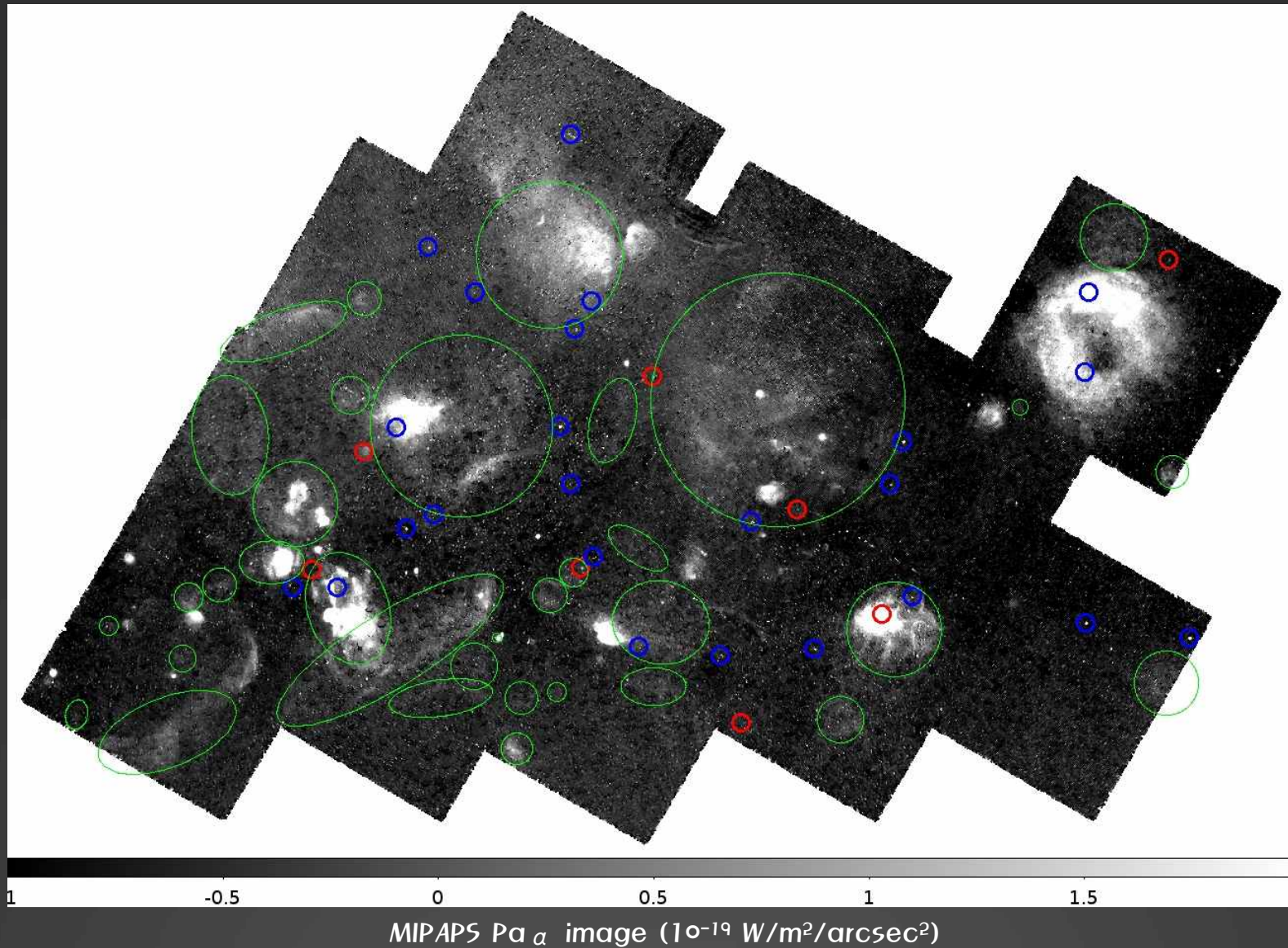
radius (arcsec), intensity (10^{-19} W/m²/arcsec²), flux (10^{-14} W/m²)

3.2

MIPAPS Pa α sources

◆ 35 large features, 32 point-like blobs

+ Point-like blobs : define 24 detections by visual inspection and 6 detections by known catalogs.



3.2

MIPAPS Pa α sources

◆ Results of visual inspection & flux photometry

Name	Positions	Radius	Overlapped with PS	WISE	CGPS	IPHAS	Paa Int	Paa Flux	Ha Int	Ha Flux	E(B-V)	Dist.	Lyc Flux	SpType
MLF01	G098.224+01.564	933	partially	Oy	Oy	Oy								
MLF02	G098.337-01.736	1830	partially	Y	Y	Y								
MLF03	G099.138+05.259	1899	partially	Oyw	N	partial, Y								
MLF04	G100.617+02.606	470	partially	Y	Y	Y	0.12	0.81	0.56	3.89	0.30	5000	47.23	>O9.5V
MLF05	G102.606-00.878	2675	partially	Oy	Oy	Oy								
MLF06	G103.450-02.294	1314	partially	Y	Y	Y	0.29	15.88	0.64	34.86	0.71	5000	48.60	O7.1V
MLF07	G104.421+02.718	7198	partially	Oy	Oy	Oy								
MLF08	G106.262-00.769	2735, 2343	partially	Oy	Oy	Oy								
MLF09	G106.376-01.772	1833, 1063	partially	YW	YW	Y								
MLF10	G106.619+00.398	1940, 844	partially	YW	YW	Y								
MLF11	G107.021+02.404	1254, 2459	partially	YW	YW	partial, Y								
MLF12	G107.621-00.000	830	partially	Y	Y	Y	0.26	5.58	0.23	4.97	1.18	5000	48.23	O8.2V
MLF13	G107.892-01.864	539	partially	Y	YW	Y	0.11	1.04	0.56	5.09	0.29	5000	47.34	>O9.5V
MLF14	G107.987-00.347	996	partially	Y	Y	Y	0.36	11.15	0.92	28.60	0.63	5000	48.43	O7.5V
MLF15	G108.014+04.993	4136	partially	Oy	Oy	partial, Y								
MLF16	G108.436-01.956	937	partially	Y	Y	partial, Y	0.04	1.20						
MLF17	G108.517-02.746	893	partially	Y	Y	Y	0.50	12.61	1.02	25.60	0.75	5000	48.51	O7.3V
MLF18	G108.822-01.019	267	partially	Oy	Y	Y	3.86	8.57	6.10	13.64	0.88	5000	48.36	O7.8V
MLF19	G109.188-01.454	1320	partially	YW	YW	Y								
MLF20	G109.397+02.308	5155	partially	Oy	Oy	partial, Y								
MLF21	G109.706-01.955	2953, 1019	partially	YW	YW	Y								
MLF22	G110.502-01.207	7407, 2090	partially	Oy	Oy	Oy								
MLF23	G110.925+04.307	957	partially	Y	Y	Y	0.15	4.38	0.48	13.70	0.52	5000	48.01	O8.7V
MLF24	G111.139+02.784	1073	partially	Y	Y	Y	0.19	6.77	0.69	24.95	0.44	5000	48.18	O8.3V
MLF25	G111.165-00.552	2291, 3242	partially	Oy	Oy	Oy								
MLF26	G112.002+01.069	2397	partially	Oy	Oy	Oy								
MLF27	G112.198+03.790	3778, 1194	partially	Y	Y	Y								
MLF28	G112.365+00.162	1833, 1210	partially	Oy	Oy	Oy								
MLF29	G113.027+02.160	2154, 3398	partially	Y	Y	Y								
MLF30	G113.177-00.202	952	partially	Y	Y	Y	0.11	3.12	0.27	7.62	0.65	5000	47.88	O9.0V
MLF31	G113.660-00.361	834	partially	Y	Y	Y	0.35	7.57	1.17	25.51	0.48	5000	48.24	O8.1V
MLF32	G113.745-01.347	769	partially	YW	YW	Y	0.10	1.80	0.16	2.92	0.87	5000	47.68	O9.3V
MLF33	G113.990-02.507	4112, 1950	partially	Y	Y	Y								
MLF34	G114.906-00.853	540	partially	Y	YW	YW								
MLF35	G115.411-02.236	620	partially	YW	Y	Y								

35 large features

32 point-like blobs

radius (arcsec), intensity (10^{-19} W/m²/arcsec²), flux (10^{-14} W/m²)

Name	Positions	Corresponding object	Types	Overlapped with PS	WISE	CGPS	IPHAS	Paa Flux	Ha Flux
MPB01	G097.980-01.026	EM* MWC645	Be	PS	PS	N	PS	1.70	1.97
MPB02	G098.265+04.907	PN K 3-60	Planetary Nebula	PS	PS	Y	PS	0.12	0.11
MPB03	G099.530+04.398	HD239712	B2Vnne	PS	PS	N	PS	0.50	
MPB04	G099.594+03.143	[KW97] 59-24	emission-line star		N	N	N	1.22	
MPB05	G099.596-00.793				N	N	N	1.11	
MPB06	G102.309-00.366	TYC3986-137-1		largely	N	N	N	0.67	
MPB07	G102.454+02.068			partially	N	N	N	1.19	
MPB08	G102.657+01.389	WR151 (V* CX Cep)	WN4+O5V	partially	PS	N	PS	0.25	0.24
MPB09	G102.783-00.653	WR153 (HD211853)	WN6+O6I	PS	PS	N	PS(Witham+08)	1.48	
MPB10	G103.853-01.184	WR154 (HD213049)	WC6	PS, partially	PS	N	PS	0.60	
MPB11	G104.111+01.001	PN Bl 2-1	Planetary Nebula	partially	PS	Y	PS	0.12	0.09
MPB12	G104.836+00.806			largely	N	N	N	1.28	
MPB13	G104.994-02.338	HD240010	B1IVnnpe (HAe/Be)	PS, largely	Y	YW	Y	0.18	2.47
MPB14	G105.323-01.294	WR155 (V* CQ Cep)	WN6+O9II-Ib	PS	PS	N	PS	1.39	0.84
MPB15	G106.391+03.093	EM* GGR71	Be (HAe/Be)	PS	PS	N	Y	0.36	0.46
MPB16	G106.601-01.148	TYC3992-728-1		largely	N	N	N	1.17	
MPB17	G107.313+00.258	TYC3996-514-1		largely	N	N	N	0.60	
MPB18	G107.341+04.278	EM* AS 492	emission-line star	PS	PS	N	only D, PS	0.66	
MPB19	G107.507+00.085	EM* GGR102	emission-line star		PS	Y	Y	0.25	0.42
MPB20	G107.605+03.835	2MASS22321530+6230118		largely	N	N	N	1.34	
MPB21	G107.673+01.400	V* V525 Cep	Be (HAe/Be)	PS	PS	N	PS	1.08	0.76
MPB22	G107.680+06.886	TYC4272-684-1		largely	N	N	no data	0.42	
MPB23	G107.844+02.315	NGC7354	Planetary Nebula		PS	Y	Y	2.73	3.28
MPB24	G109.177+04.414	BHB; 2MASS22414172+6344505		largely	N	N	N	0.36	
MPB25	G109.824+00.919	WR156	WN8	PS, partially	PS	N	only D, PS	1.36	
MPB26	G109.916+05.112	2MASS22443929+6443220			N	N	N	0.32	
MPB27	G110.255+00.702			partially	N	N	N	1.13	
MPB28	G110.407+02.294	IRAS22577+6209; 2MASS22594539+622		largely	Y	YW	N	0.65	
MPB29	G110.905+01.897	EM* AS 505	B4e (HAe/Be)	PS	Y	N	PS	0.47	
MPB30	G111.334-00.236	WR157 (HD2194608)	WN4.5+B1II	PS	PS	N	PS	0.68	
MPB31	G111.730+00.035	EM* MWC1080	B0eq (HAe/Be, H)	PS	PS	N	Y	2.21	1.32
MPB32	G112.026-00.242	2MASS23203291+6041084		largely	N	N	N	0.55	

3.3 Scientific potential of MIPAPS Pa α data

◆ WISE H II region sources

+ 49 **Known** sources.

- MIPAPS Pa α detections : 42, IPHAS H α detection : 44.
- Only IPHAS H α detection : 5 (overlapped with other sources).

+ 50 **Candidate** sources.

- MIPAPS Pa α detections : 25, IPHAS H α detection : 26.
- Only MIPAPS Pa α detection : 5 (high extinction).
- Only IPHAS H α detection : 7 (small size, overlapped with other sources).

+ 21 **Group** sources.

- MIPAPS Pa α detections : 16, IPHAS H α detection : 18.
- Only IPHAS H α detection : 4 (small size, overlapped with other sources).

+ 92 **Radio quiet** sources.

- MIPAPS Pa α detections : 6, IPHAS H α detection : 23.
- Only MIPAPS Pa α detection : 1 (high extinction).
- Only IPHAS H α detection : 18 (small size, overlapped with other sources).

→ MIPAPS Pa α data are expected to identify >1000 candidates as H II regions for the whole plane.

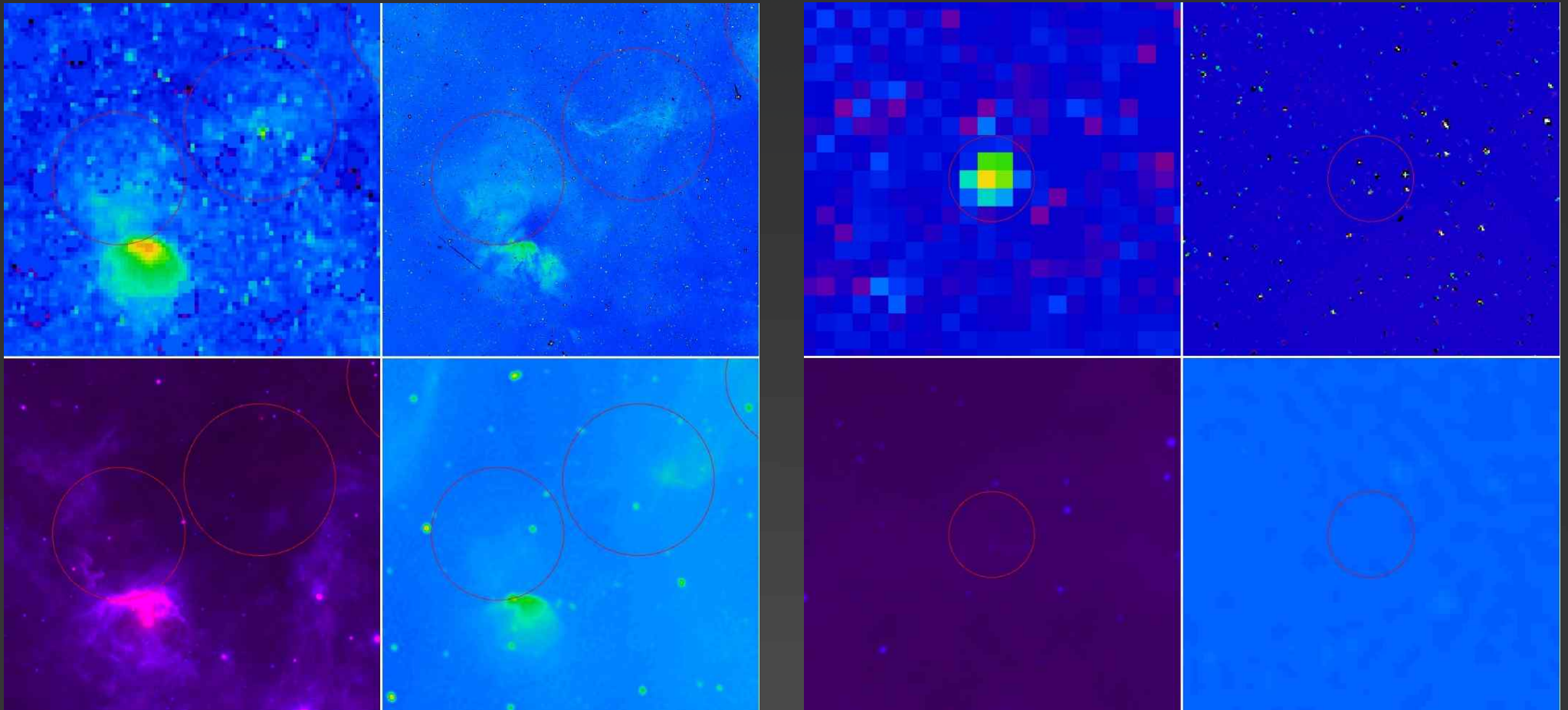
3.3

Scientific potential of MIPAPS Pa α data

◆ MIPAPS Pa α sources

+ 35 Large features : 13 have no corresponding known sources.

+ 32 Point-like blobs : 14 no counterparts, 3 PNe, 6 WRs, 9 emission-line stars (including 5 Herbig).



MLF30 & MLF31

MPBo5

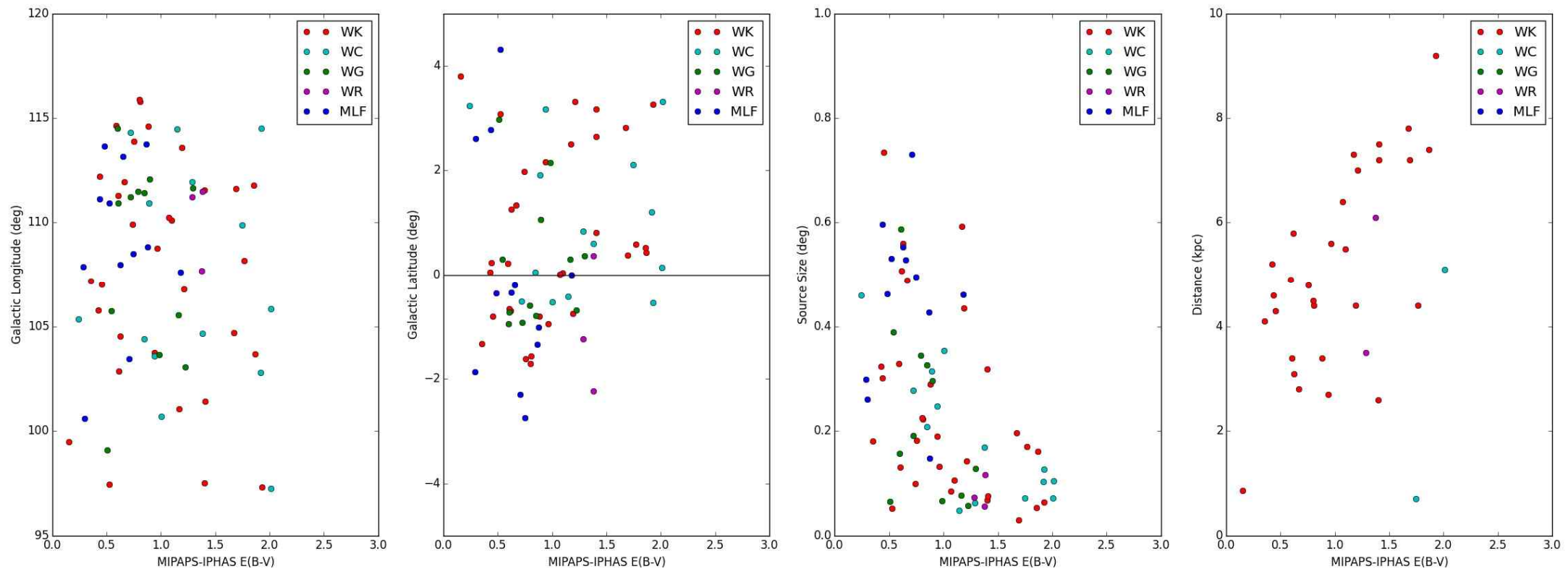
Upper : MIPAPS Pa α , IPHAS H α

Lower : WISE 12 μ m (blue) + 22 μ m (red), CGPS radio continuum

3.3 Estimations of dust extinction, distance, spectral type

◆ MIPAPS-IPHAS E(B-V)

- + Radiative recombination of Hydrogen (from Draine, Physics of the interstellar and intergalactic medium).
- + Assuming Case B & $T=10^4$ K (general H II region), **observed $\text{Pa}_\alpha / \text{H}_\alpha$** values give **E(B-V)**.

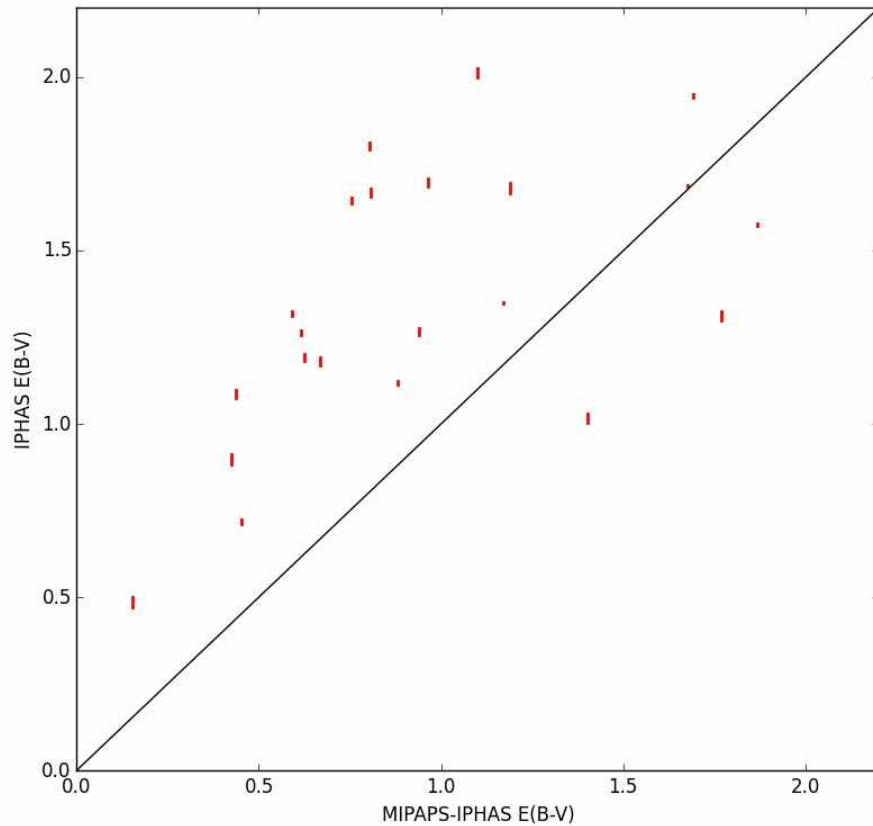


MIPAPS-IPHAS E(B-V) **vs.** source position (l, b), size, distance

3.3 Estimations of dust extinction, distance, spectral type

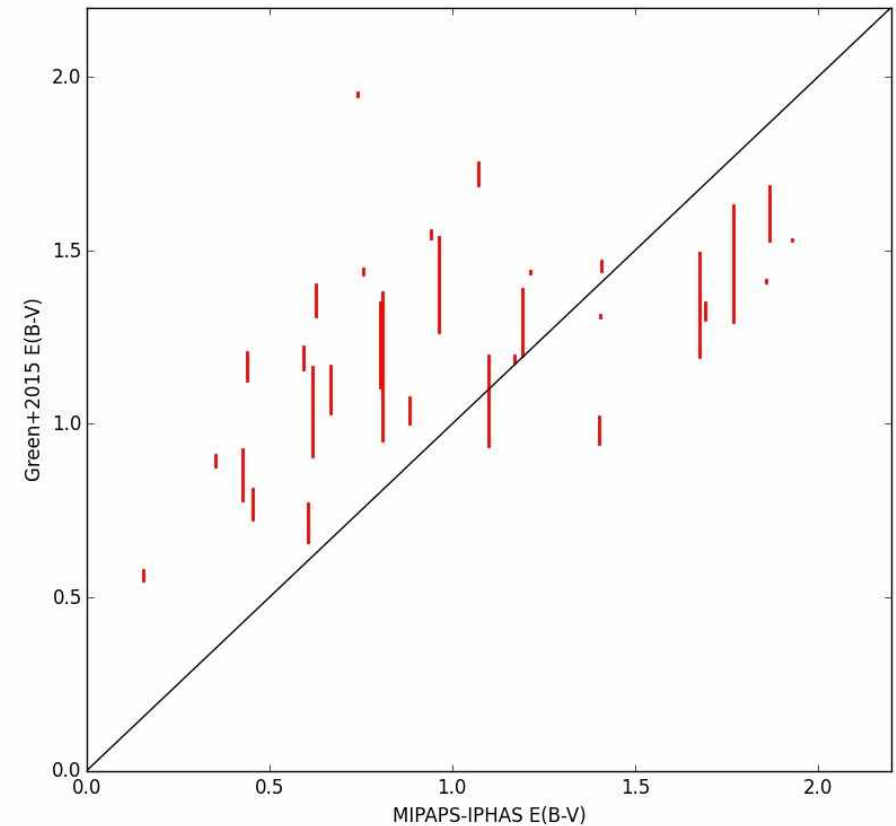
◆ MIPAPS-IPHAS E(B-V) (for WISE Known sources)

+ Comparison with other 3D dust extinction data.



MIPAPS-IPHAS E(B-V)

vs. IPHAS E(B-V)



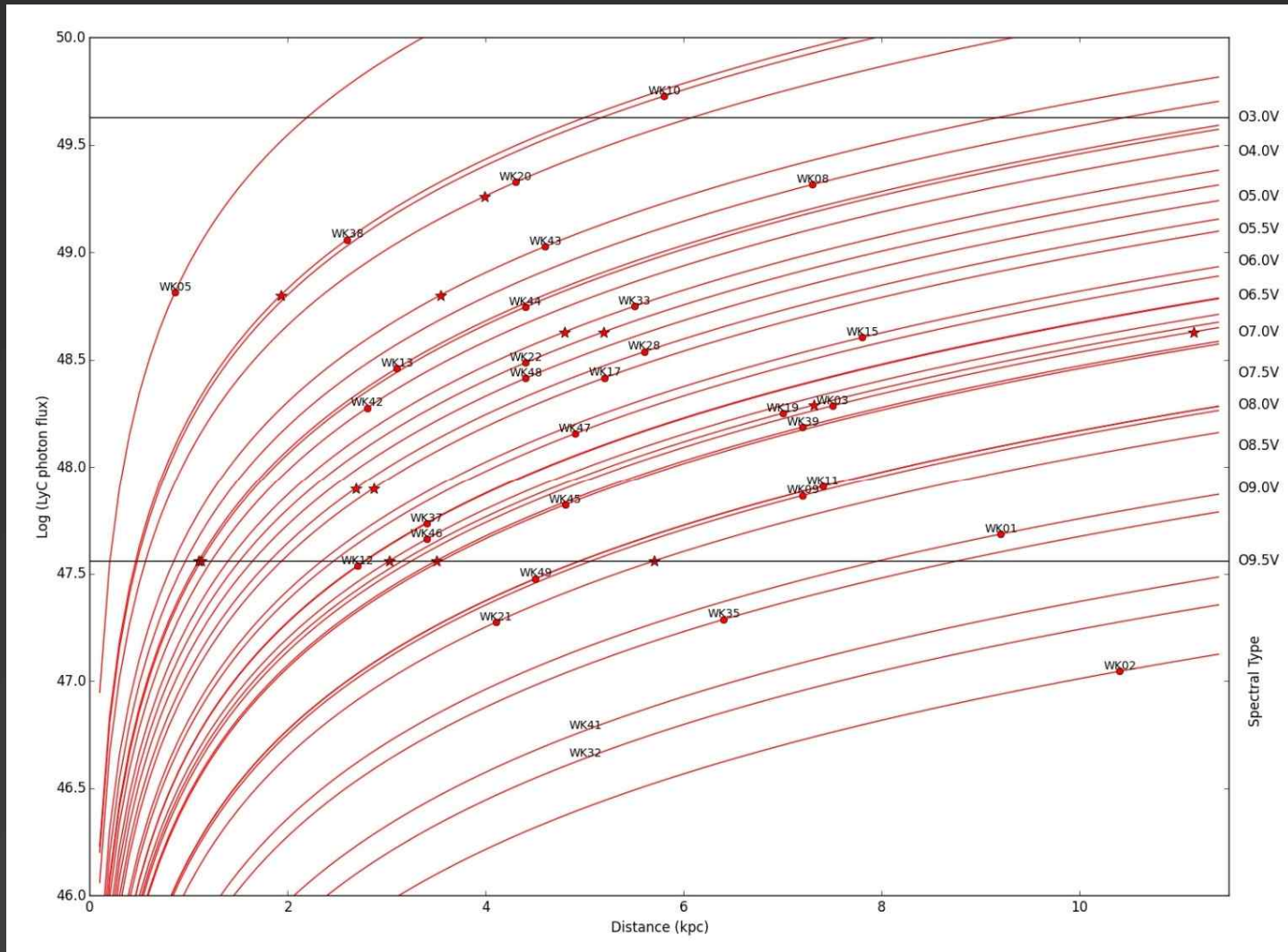
vs. Pan-STARRS1-2MASS E(B-V) (Green+ 2015)

3.3

Estimations of dust extinction, distance, spectral type

◆ Distance-Spectral type of ionizing source (for WISE Known sources)

- + If the **distance** is known, intrinsic total flux of Pa_α (or H_α) can be calculated. (\rightarrow LyC photon flux)
- + Lyman continuum photon fluxes as a function of **spectral types** (Martins+ 2005).



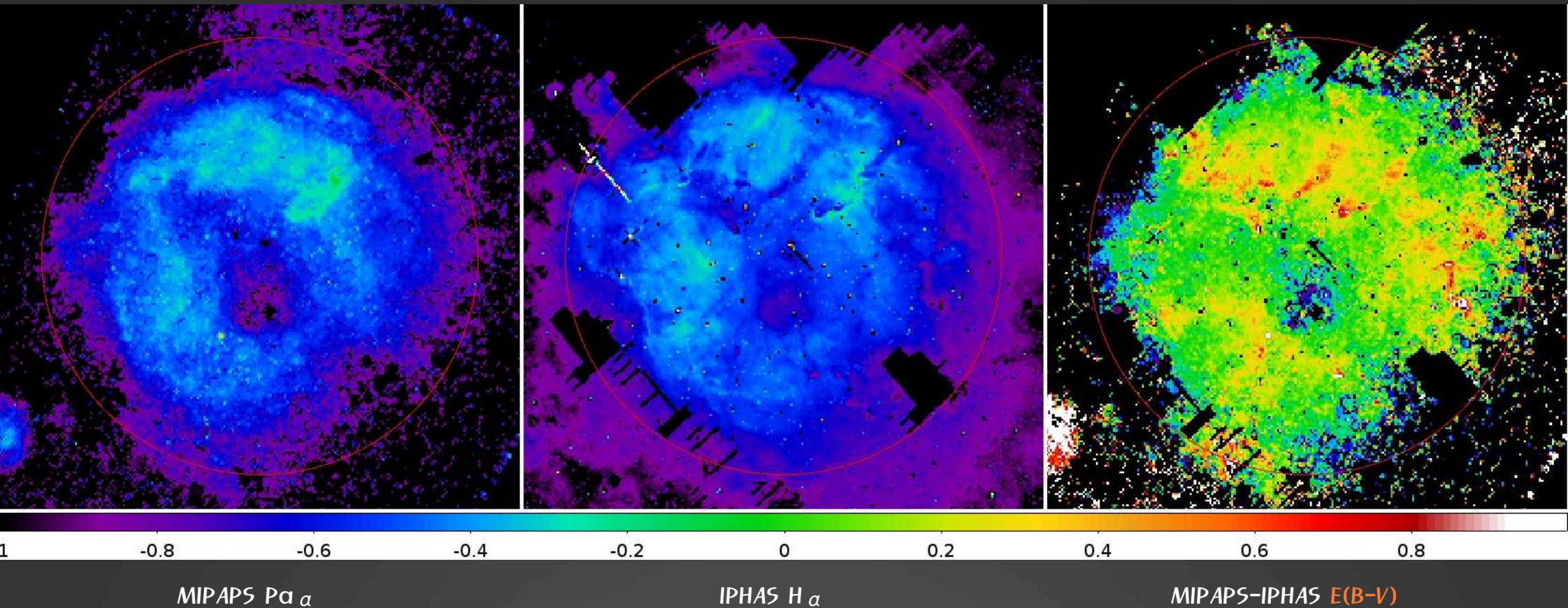
Distance vs. Estimated spectral type of ionizing sources

3.3

E(B-V) maps from Pa_α to H_α ratio

◆ WKo5 (Sh2-131)

- + Clumpy **high Pa_α** regions : due to dust extinction by foreground local clouds.
- + Diffuse **high H_α** regions : due to **dust scattering**.

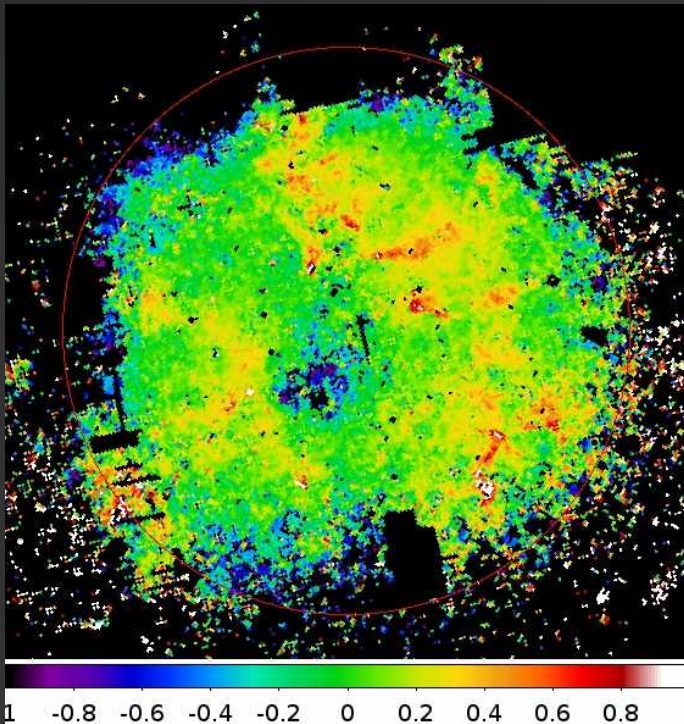


3.3

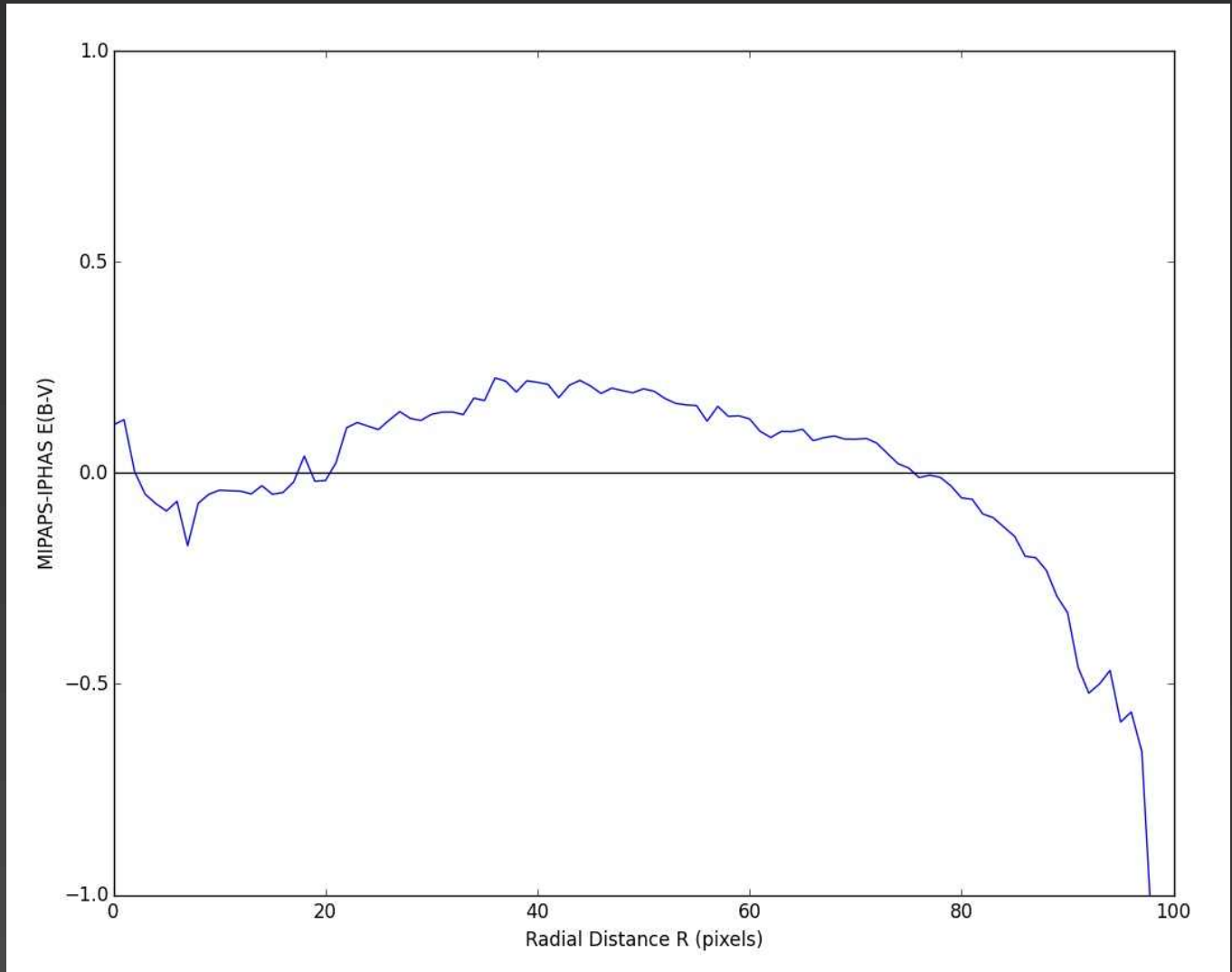
$E(B-V)$ maps from $\text{Pa } \alpha$ to $\text{H } \alpha$ ratio

◆ WkO5 (Sh2-131)

+ Plot $E(B-V)$ radial profile from the center.



$E(B-V)$ map for WkO5



Radial profile of $E(B-V)$

Next plan

◆ Data release 1 (June, 2017)

- + Cepheus region (Q2) : comparison with WISE, IPHAS.
- + Carina region (Q4) : comparison with WISE, SHS.

◆ Data release 2 : all data including edge-shadowing correction.

- + The *whole* Galactic plane : MIPAPS Pa α source catalog.